

KO ANR COMPASS

GEPI

Coordinateur: M. Puech

Roadmap instrumentale E-ELT

Instruments de 1ère lumière :

- ELT-CAM
- ELT-IFU

Chaque instrument à un espace des paramètres propres distinct des autres qui requière des simulations spécifiques

Principe :

Explorer ces espaces en suivant la Roadmap E-ELT et en « priorétisant » selon les participations françaises

Year	ELT-IFU	ELT-CAM	ELT-MIR	ELT-4 (MOS or HIRES)	ELT-5 (MOS or HIRES)	ELT-6	ELT-PCS
2012	Decide science requirements, AO architecture.		VISIR start on-sky	Develop science requirements for MOS/HIRES			Call for proposals for ETD
2013			TRL Review	Call for proposals for MOS/HIRES			
2014							
2015				Selection ELT-MOS/HIRES		Call for proposals	
2016							
2017							TRL check
2018							TRL check
2019						Selection	TRL check
2020							TRL check
2021							TRL check
2022 Tel technical first light							
	Pre-studies taking the form of phase A or delta-phase A work and/or ESO-funded Enabling Technology Development (ETD)						
	Decision point						
	Development of Technical Specifications, Statement of Work, Agreement, Instrument Start.						

Definition (& Performances) of E-ELT instruments

Keeping in mind that when we talk about the VLT (They are 4 telescopes)

→ More than 12 instruments

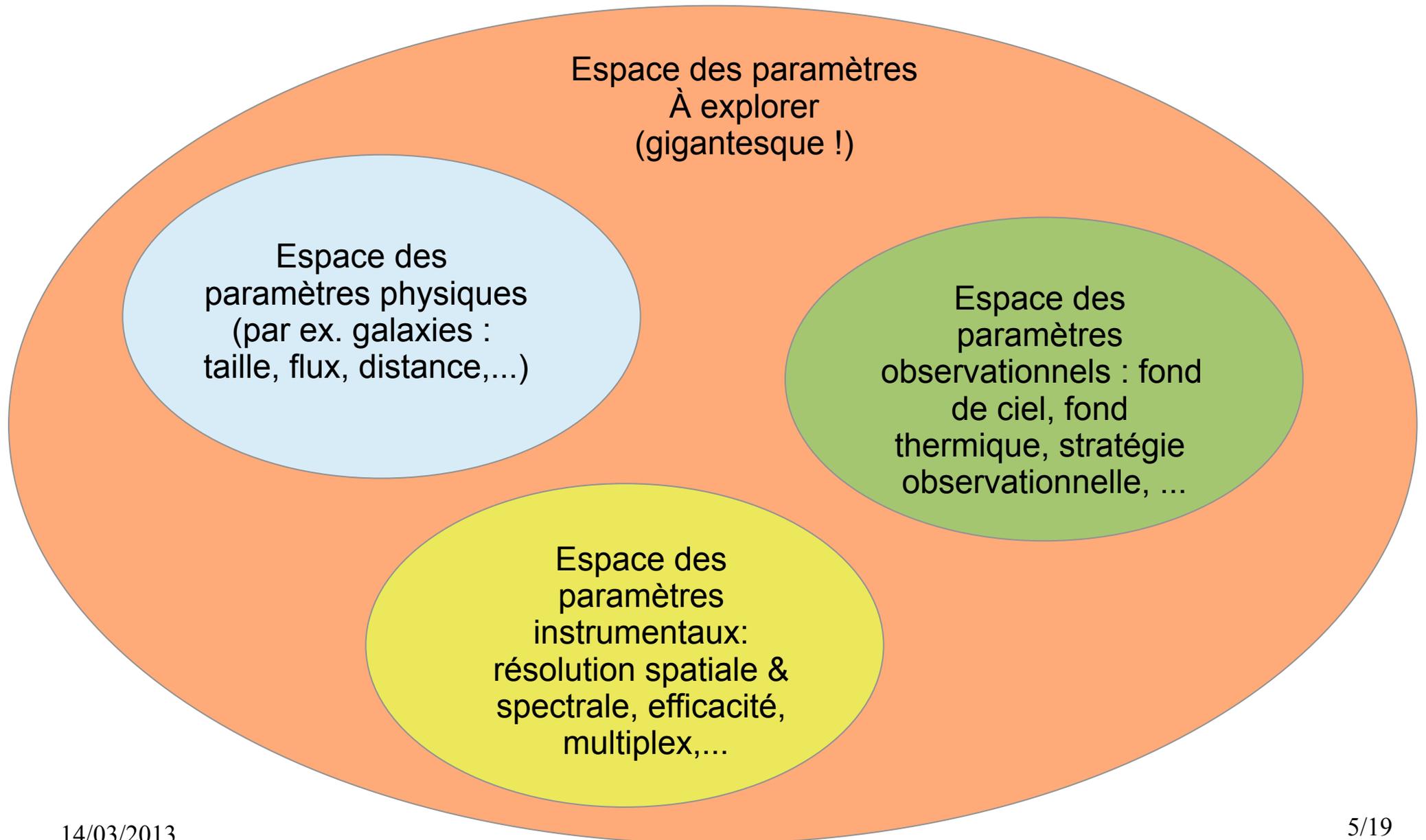
The E-ELT will be unique with few instruments

We must maximize the scientific return

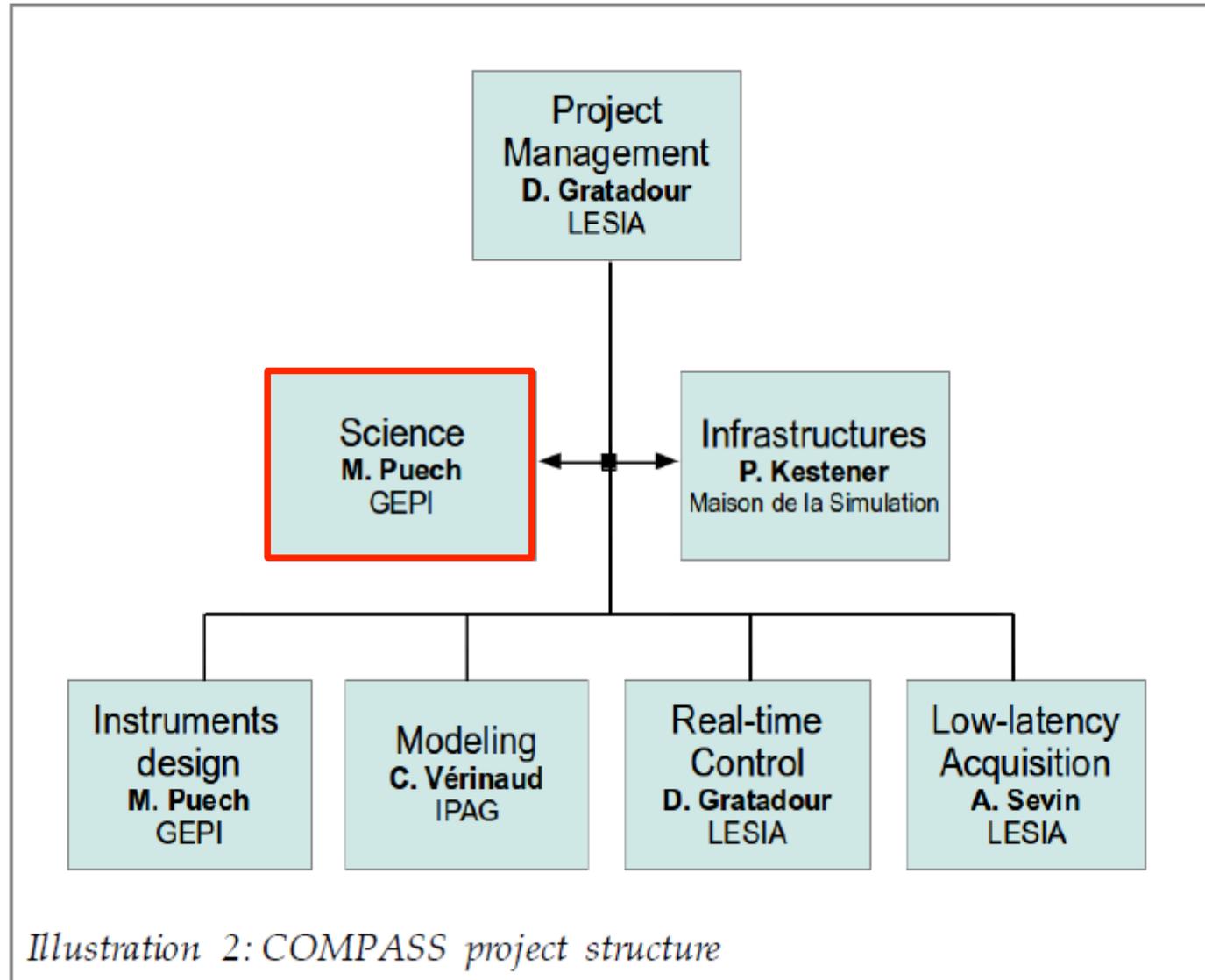
Priorités françaises

- 1ère lumière :
 - ELT-CAM : Participation GEPI-LESIA
 - ELT-IFU : Participation LAM ?
- 1ère génération :
 - ELT-MOS : consortium international MOSAIC
(France : GEPI-LESIA-LAM)
- COMPASS a un rôle majeur à jouer dans la simulation de ces instruments
- Planning ESO : décalage de 6-12 mois ?

Difficulté

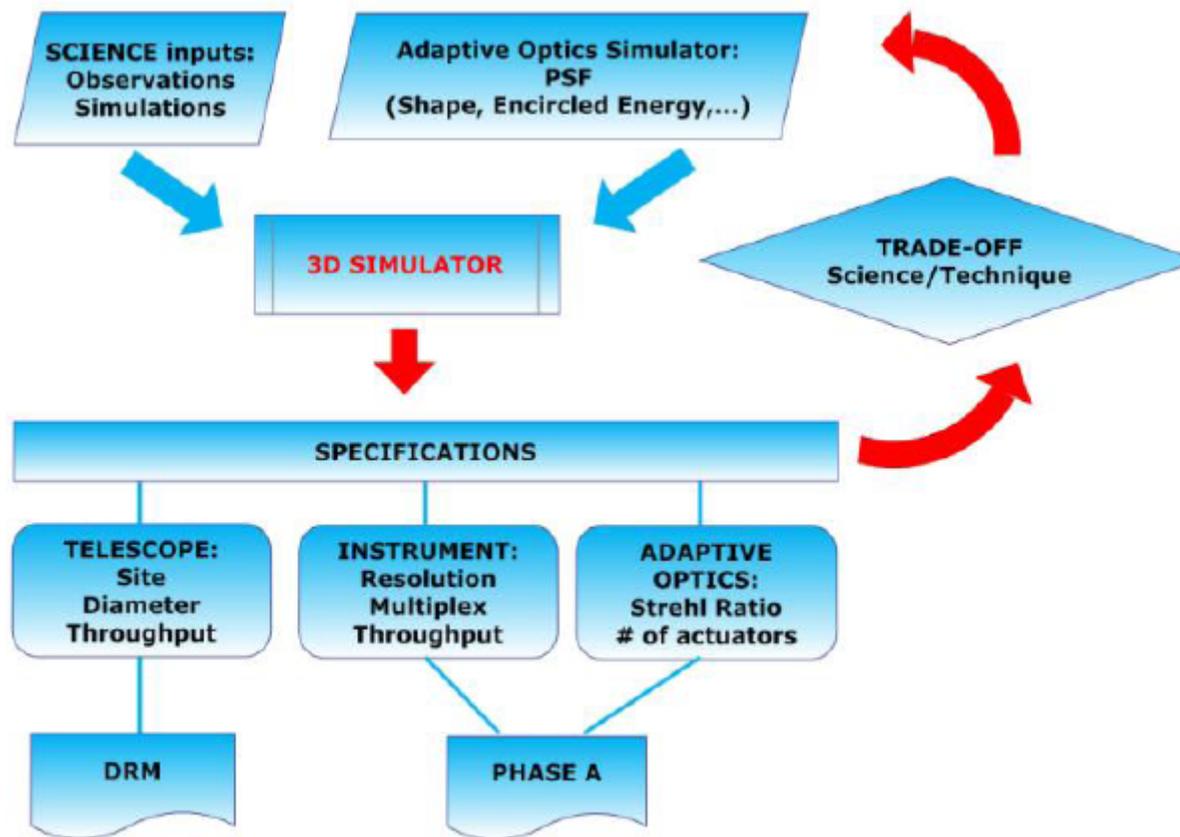


Rôle du GEPI

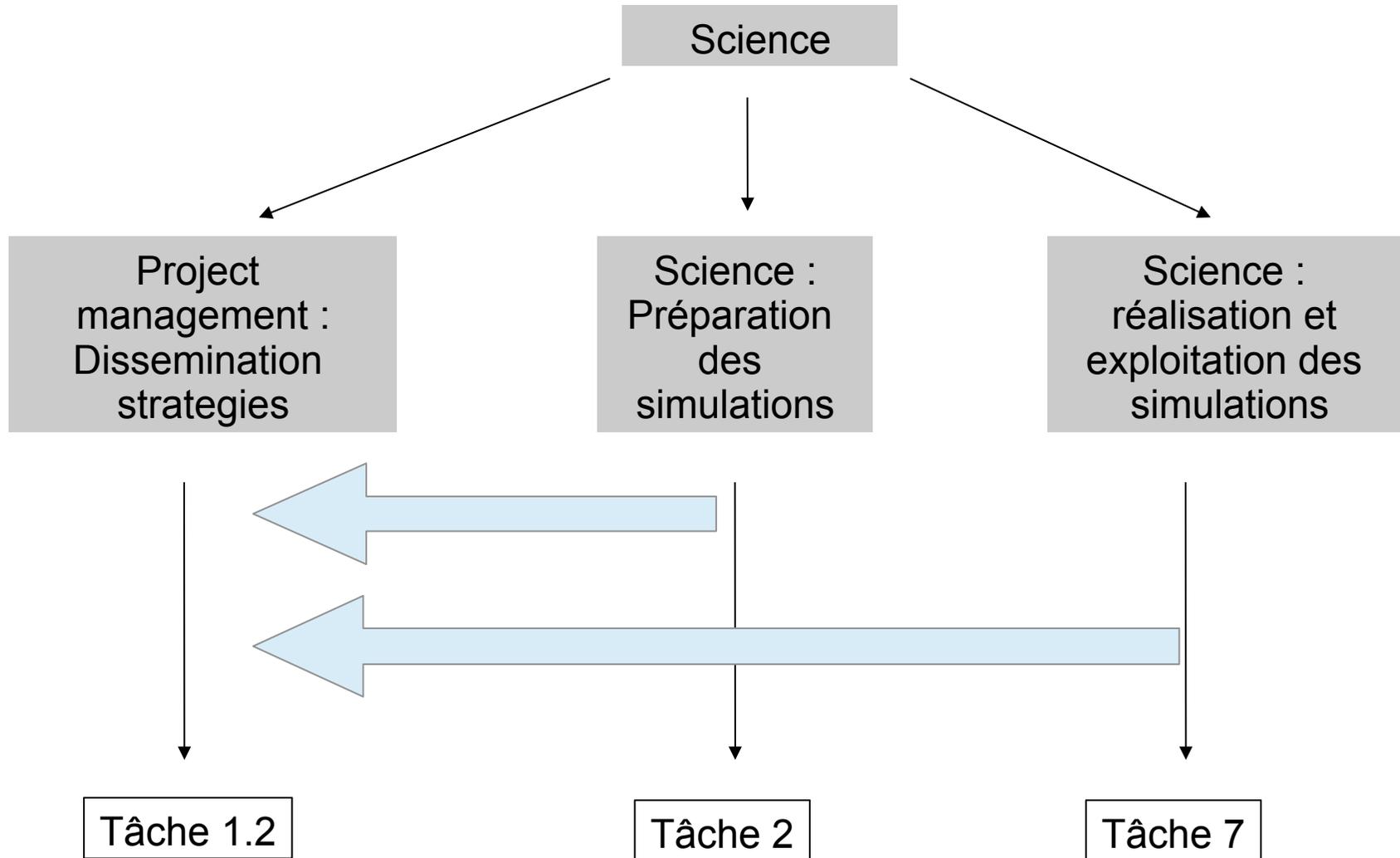


Objectifs scientifiques

Mise en place une plate-forme de simulation instrumentale :



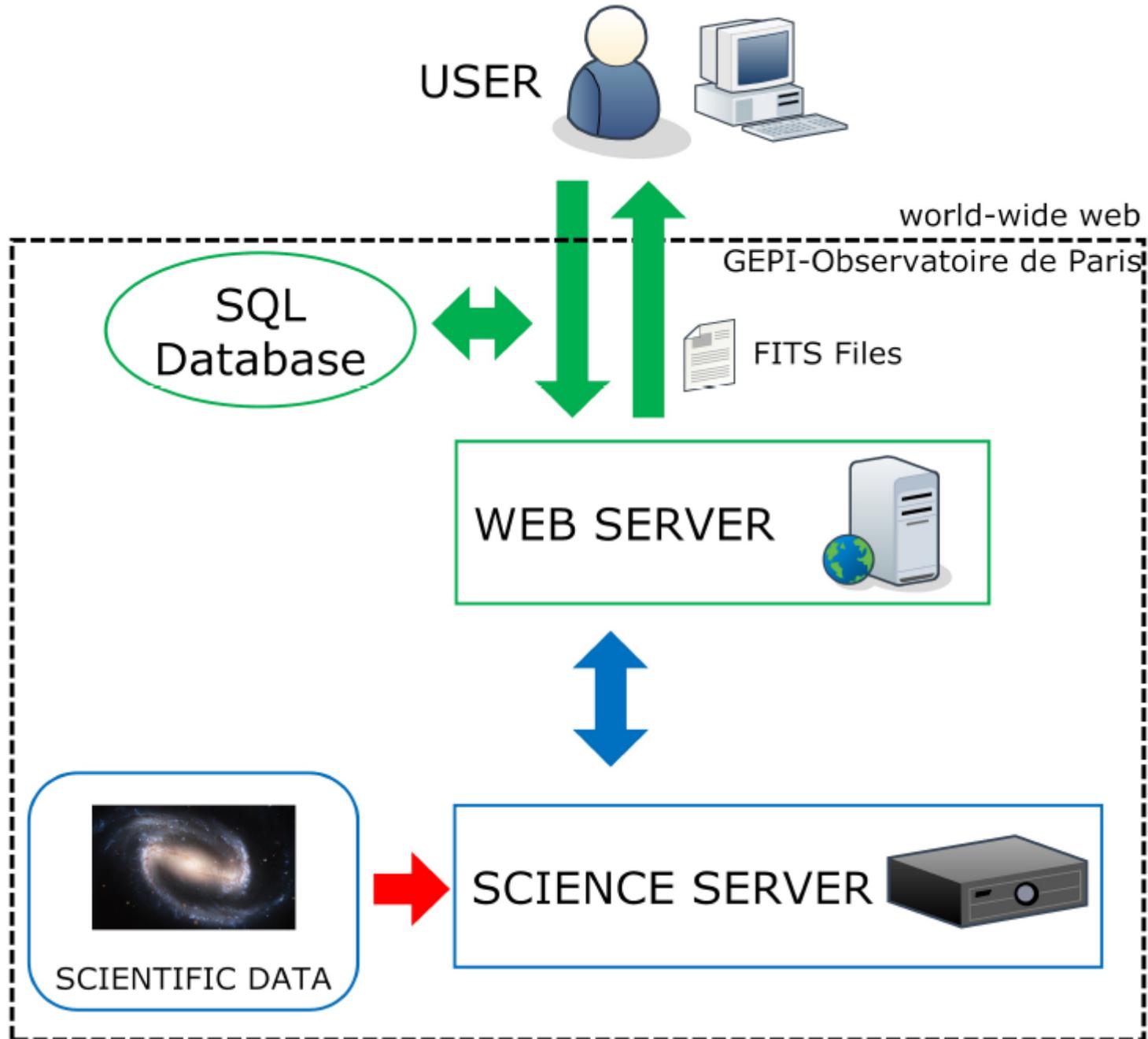
Tâches sous la responsabilité du GEPI



Tâche 1.2 : Mise en place des outils de simulation

Task 1.2 Dissemination strategies	Resp: M. Puech Total Staff : 8m.m;	Duration: 24m Temp. : 1 m.m
<p>Objectives: Define and implement dissemination strategies</p> <p>Contributors: M. Puech (2m.m), Y. Yang (1 m.m), S. Taburet (2m.m), I. Jégouzo (1m.m), T. Fenouillet (2m.m)</p> <p>Content, Milestones and Outputs:</p> <ul style="list-style-type: none">- Coordinate the interactions between <i>Science</i> and <i>Instrument design</i> tasks- Organize regular progress meetings (every 2 months)- Setting up and managing content for the AO PSF library- Setting up and managing a web-based interface for running end-to-end instrument simulations for ELT-CAM, ELT-IFU, and ELT-MOS- T0 + 9m: Preliminary version of the web-based interface- T0 + 24m : D1.2 – final public release of the web-based interface + AO PSF library		

Prototypes Phase A



Besoins matériels

- Un ordinateur scientifique multi-CPU pour répondre à la demande « nationale » : serveur Dell PowerEdge R810 (10k€)
- Serveurs web permettant un accès sécurisé et fiable au simulateur : 2xDELL Poweredge R410 (5k€ total, un sur Paris, un sur Meudon)
- Développement et « publicité » du simulateur dans les conférences : 2xMacBook Pro x2 (total 4k€)

Tâche 2.1 : définition des paramètres à explorer et de la stratégie d'exploration

TASK 2 – SCIENCE		RESP: M. Puech
Task 2.1 Interfaces to instruments models	Resp: M. Puech Total Staff: 6m.m;	Duration: 2m Temp. : 2m.m
<p>Objectives: Define interfaces between instrument and AO simulation codes</p> <p>Contributors: M. Puech (1m.m), H. Flores (1m.m), K. Disseau (1m.m), D. Mouillet (1m.m), Post-doc GEPI (2m.m)</p> <p>Inputs and/or relation to other tasks: Inputs from task 1.3 (Science req. document) and predecessor of task 2.2 <i>Preparatory studies</i></p> <p>Content, Milestones and Outputs:</p> <ul style="list-style-type: none"> - Split the simulation parameter space into instrument and AO simulation codes - Define the content of the PSF library and associated parameter space - Define the web-base interface for running end-to-end instrument simulations - Define hardware communications between the web server and the science server - T0 + 2m : D2.1 - Interface document 		

Prototypes Phase A

Welcome **Mathieu Puech** Enter expert mode logout

E-ELT/EAGLE Simulator

NOTE: when using results from the simulator, please refer to [Puech et al. \(2008\)](#)

Telescope

- E-ELT:
 - $M_1 = 42$ m (Diameter)
 - Effective central obscuration = 9% of the total collecting area.
- Total throughput = **0.216** (=Telescope&Atmosphere * Instruments * Detectors)

Instrument

IFU

- IFU size × arcsec²
- Slicer width = 37.5 mas/pixel
- Spectral Resolution R = 4000

CCD

- dark = 0.01 electrons/pixel/second
- read-out noise = 5 electrons/pixel

Target

Rotating flat ellipse

Emission line galaxy (single line)

- Template:
- Redshift z =
- Continuum Magnitude = (AB)
- Emission line =
- EW = Å (rest frame)
- $V_{max} =$ km/s
- Galaxy Diameter = arcsec

AGN case

Star case

External cube

Observation conditions

- Seeing:
- Exposure time = × s (Saturation of the detector is not taken into account.)
- Thermal background: Temperature= 280,240,150 K, Emissivities= 0.15,0.15,0.69
-

Submit simulation Reset

Welcome **Hector Flores** (The system is under testing.) Logout

E-ELT/EVE Simulator

NOTE: when using results from the simulator, please refer to [Puech et al. \(2008\)](#)

Telescope

- E-ELT:
 - $M_1 = 42$ m (Diameter)
 - Effective central obscuration = 9% of the total collecting area.
- Total throughput = **0.9** × **0.3** × **0.8** = **0.216** (=Telescope&Atmosphere * Instruments * Detectors)

Instrument

IFU

IFU MONO Mini Large

size: object IFU IFU*

" " × " " × "

Pixel size = arcsec/pixel

Spectral arm

- Spectral resolution and arm:
- Spectral sampling = pixels per spectral resolution.
- Lambda range = central lambda +/- 750 Å

CCD

Optical CCD Infrared CCD

dark = electrons/pixel/second dark = electrons/pixel/second

read-out noise = electrons/pixel read-out noise = electrons/pixel

Target

Rotating flat ellipse

- Position Angle = degrees
- Major axis = arcsec
- Minor axis = arcsec
- H α EW = Å (rest frame)
- Continuum magnitude = (AB)
- Redshift z =

Halo case

High z galaxy

Stellar case

Planet case

IGM case

Observation conditions

- Seeing:
- Exposure time = × s (Saturation of the detector is not taken into account.)
- Thermal background:
 - Temperature = K
 - Emissivities =
- Include DAR (differential atmosphere refraction) at zenith distance degrees
- Mauna Kea sky background

Submit simulation Reset

* Under development.

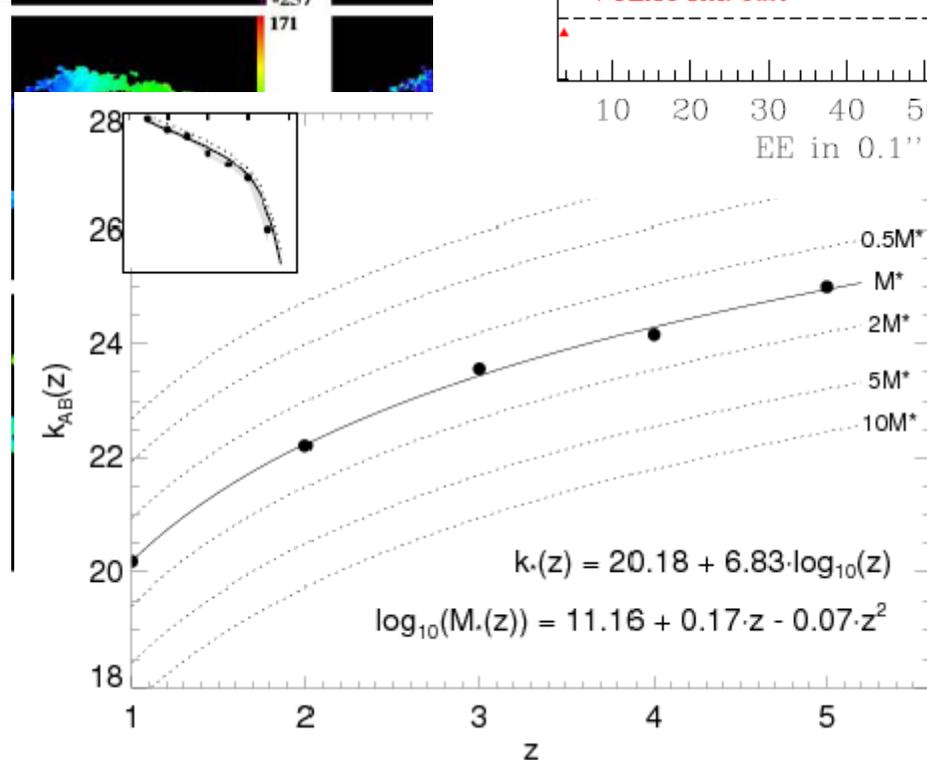
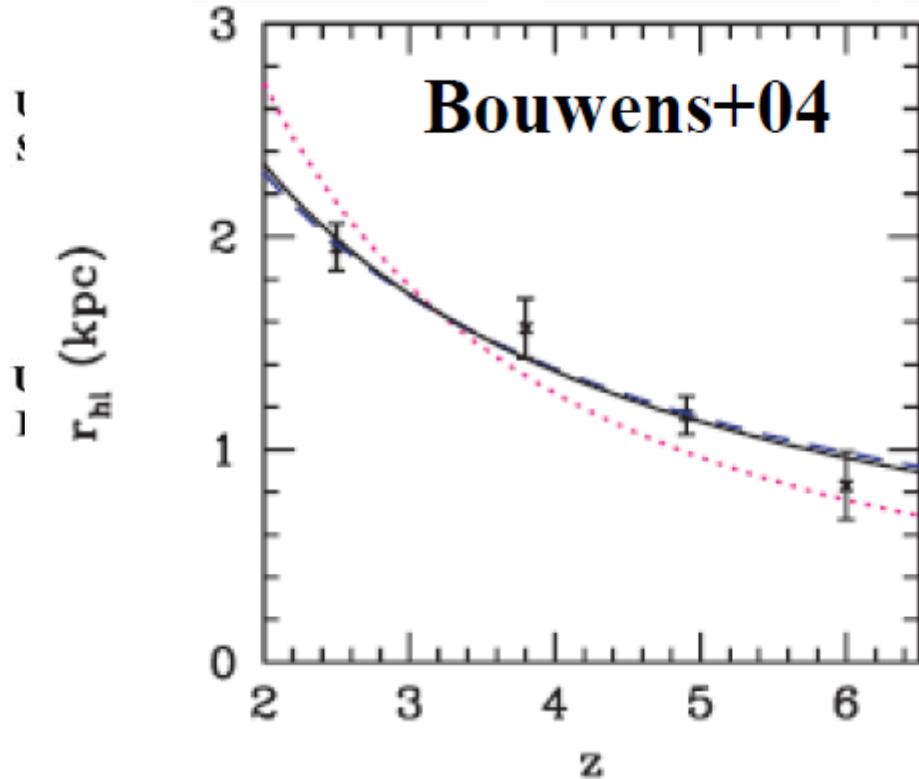
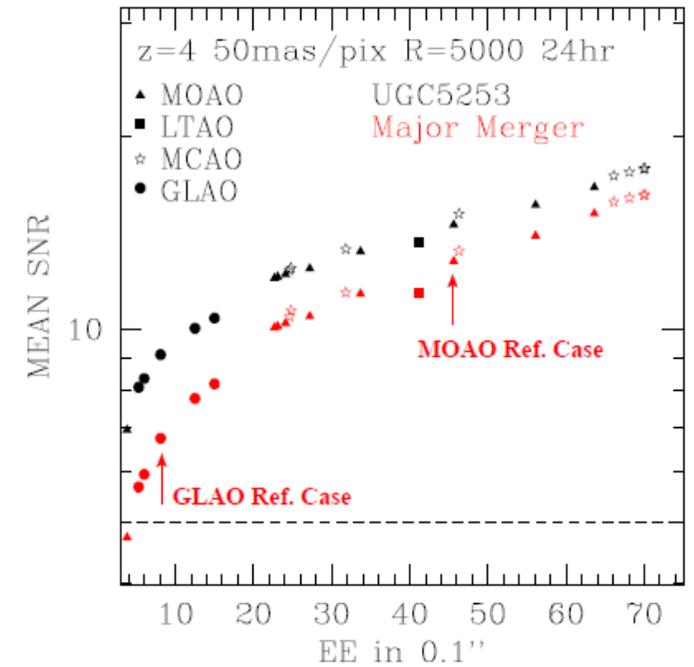
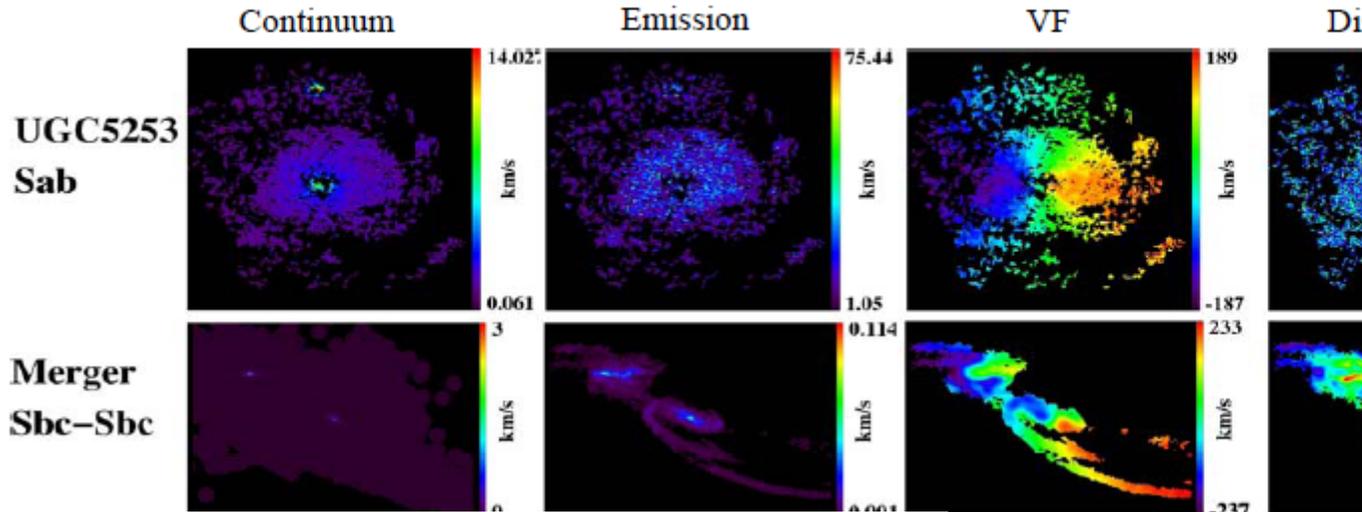
E-ELT/EVE Simulator: The Web Interface v0.5

Prototypes préliminaires: manquent plusieurs aspect dans la simulation
Ciel, absorption telluriques etc.
et une base de données des galaxies proches.

Tâche 2.2 : préparation et test des données d'entrée des simulations

Task 2.2 Preparatory studies	Resp: M. Puech Total Staff: 17m.m;	Duration: 4m Temp. : 8m.m
<p>Objectives: Set up and first tests of software solutions for instrument simulations</p> <p>Contributors: M. Puech (2m.m), F. Hammer (2m.m), H. Flores (2m.m), K. Disseau (2m.m), Y. Yang (2m.m), S. Taburet (1m.m), I. Jégouzo (2m.m), Post-doc GEPI (4m.m)</p> <p>Inputs and/or relation to other tasks: Inputs from task 2.1 (<i>interfaces to instruments models</i>). Predecessor of task 7 <i>instruments design studies</i>.</p> <p>Content, Milestones and Outputs:</p> <ul style="list-style-type: none">- T0+3m: Set-up and tests of hardware simulation tools. Define the astrophysical input data needed to run task 7- T0+3m: Design of software solutions (public web site, web-based interface for instrument simulations, PSF library)- T0+4m: Prepare the astrophysical data needed to run task 7- T0+5m: Implementation of a preliminary version of the instrument simulation code. Implement the astrophysical data needed to run task 7- T0+6m: Firsts tests of instrument simulations using simple astrophysical cases reproducing observations in the literature. Definition of scientific figures of merit for comparing simulated data of real observed cases in the literature. Science assessment of the instrument simulations		

Exemple « galaxies »



CALIFA

(Sanchez et al., 2012)

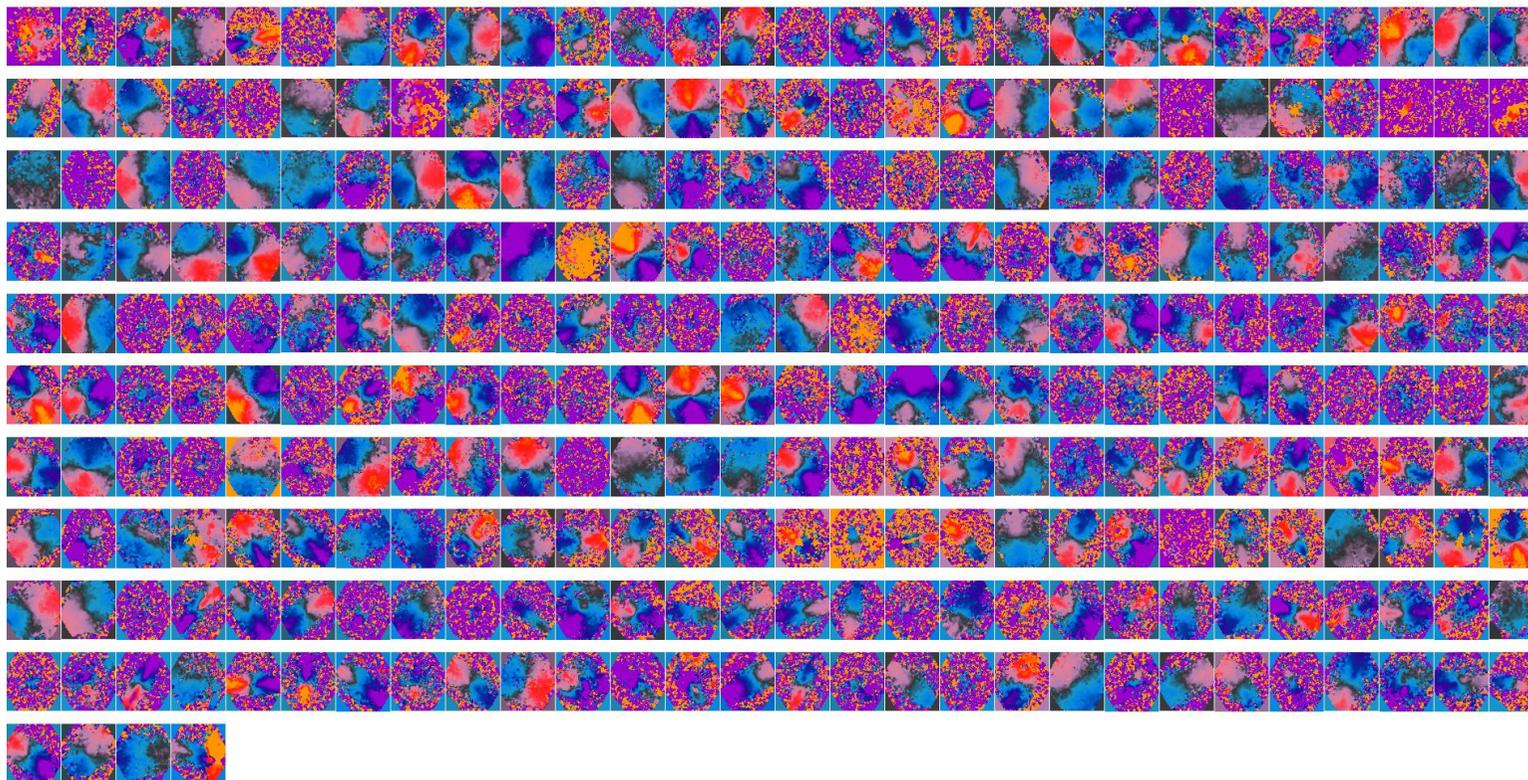
more than 600 galaxies selected on the SDSS survey

IFU: PMAS/PPAK at calar alto. 2D maps of :

(a) stellar populations: ages and metallicities;

(b) ionized gas: distribution, excitation mechanism and chemical abundances; and

(c) kinematic properties: stellar and ionized gas.

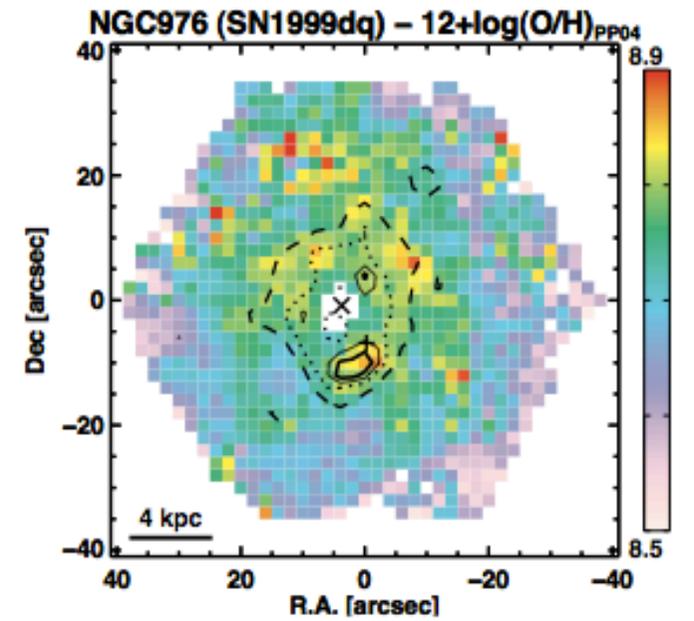
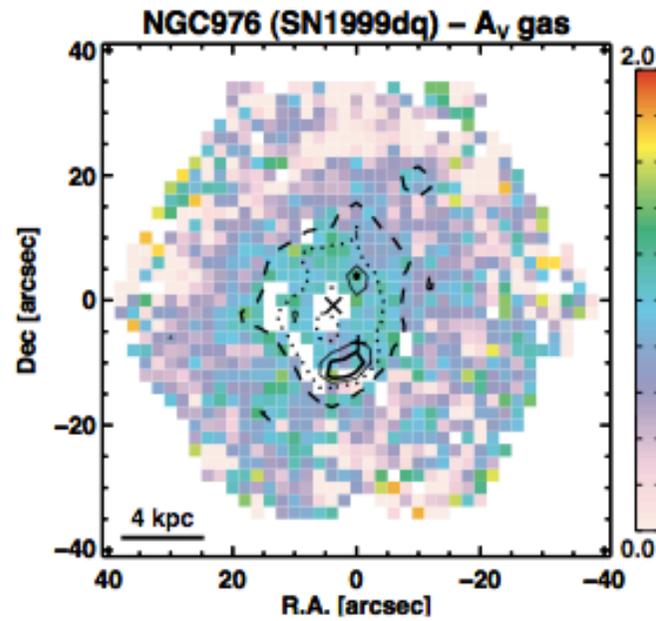
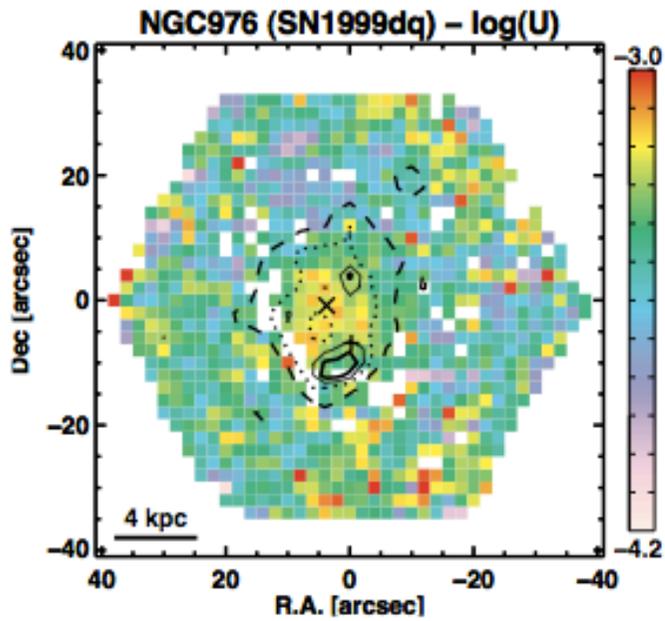
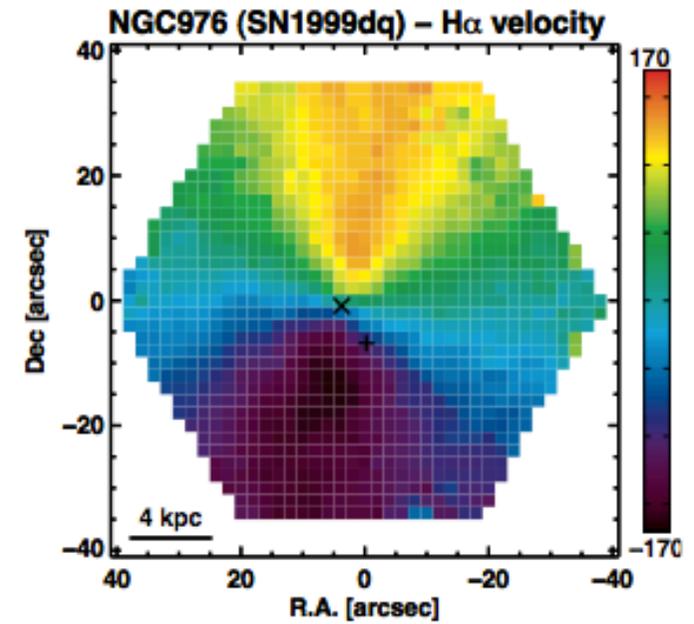
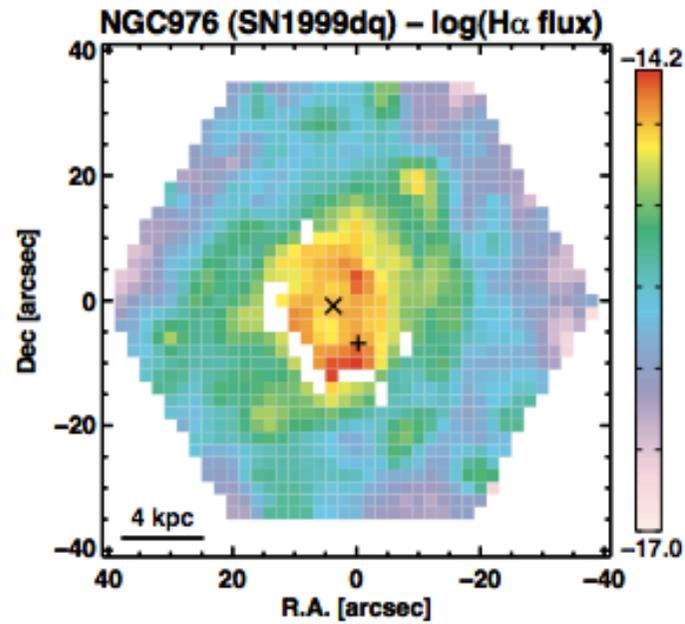


284 velocity field



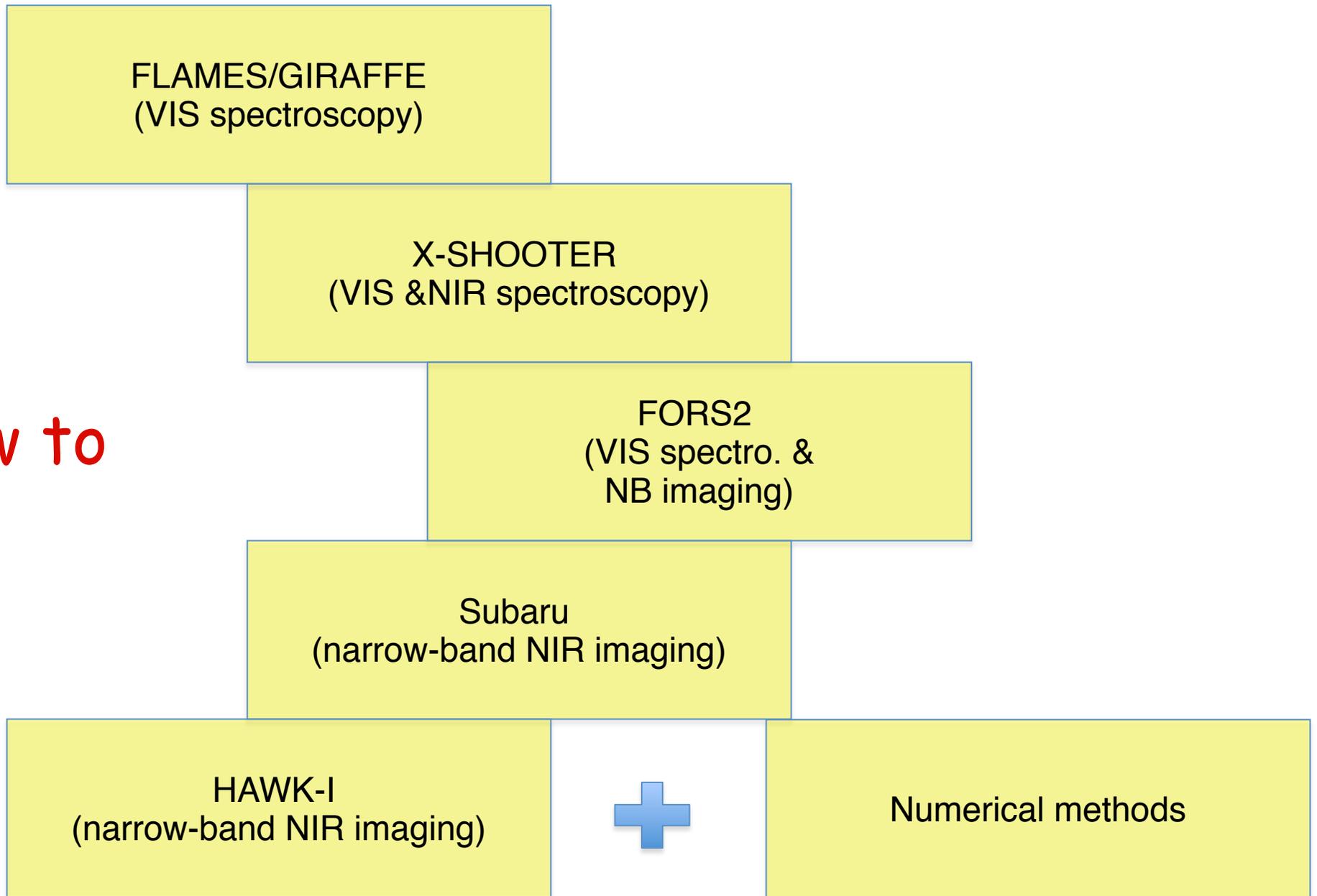
MOSAIC





To characterize the sky background: as usual need very good data ...

How to



Should provide spatial variations up to a few arcmin² @ 0.9-1.0 μ m

Conclusion of the ongoing work

FLAMES/GIRAFFE
(VIS spectroscopy)

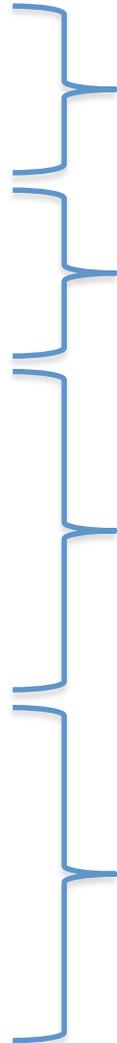
X-SHOOTER
(VIS & NIR spectroscopy)

FORS2
(VIS spectro.)

Numerical methods

FORS2 & HAWK-I
(narrow-band)

Subaru
(narrow-band NIR imaging)



Rodrigues et al, 2012

J. Vernet work

Yang et al. 2012



Puech et al.2012 & K. Disseau

And the telluric lines?

The method (one of them Vacca et al 2003)
we just need to include the telluric line in the
simulator to study the effect on the S/N.

Tâche 7.2 : Mise en œuvre et exploitation des simulations

TASK 7 – E-ELT INSTRUMENTS DESIGN

RESP: M. Puech

Task 7.2 ELT-MOS / ELT-IFU

Resp: Post-doc GEPI

Duration: 12m

Total Staff: 23m.m;

Temp.: 12m.m

Objectives: End-to-end simulations of IFU data for the E-ELT. Exploration of the science, observational and instrument parameter space. Impact on the instrument design.

Contributors: G. Rousset (1m.m), Post-doc GEPI (12m.m), M. Puech (2m.m), H. Flores (2m.m), F. Hammer (2m.m), K. Disseau (2m.m), Y. Yang (2m.m)

Inputs and/or relation to other tasks: Inputs from task 1.3 (*Science requirements*) and task 2.2 (*preparatory studies*)

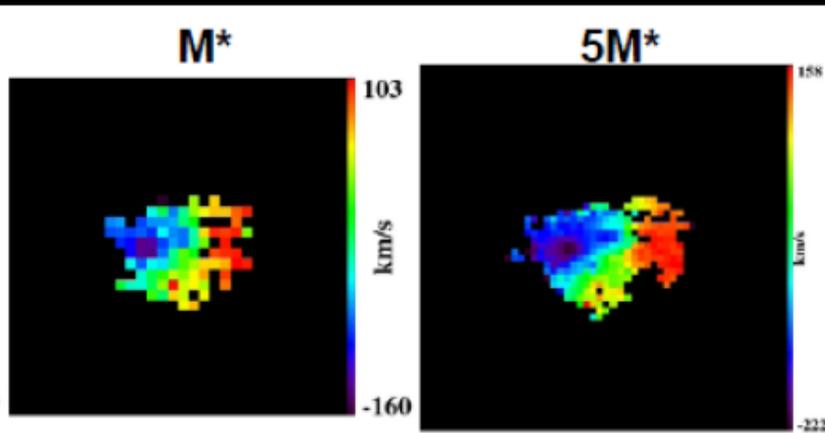
Content, Milestones and Outputs:

- T0+13m: Definition of the parameter spaces
- T0+14m: ELT-IFU Simulation run. Library of 3D fake data for ELT-IFU.
- T0+17m: Analysis of ELT-IFU simulations. Impact on ELT-IFU design.
- T0+19m: ELT-MOS Simulation run. Library of 3D fake data for ELT-MOS.
- T0+22m: Analysis of ELT-MOS simulations. Impact on ELT-MOS design.
- T0+24m: Final report (D7.2) and scientific paper

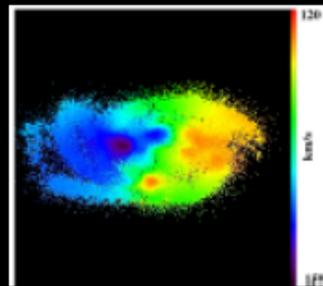
Exemple (1) : E-ELT DRM

GLAO leads to smaller S/N compared with MOAO and will limit observations to smaller-mass galaxies. GLAO will impact strongly the recovery of Rotation Curves and detailed kinematics.

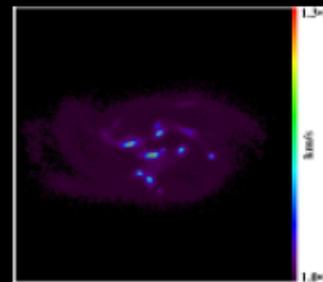
MOAO z=4



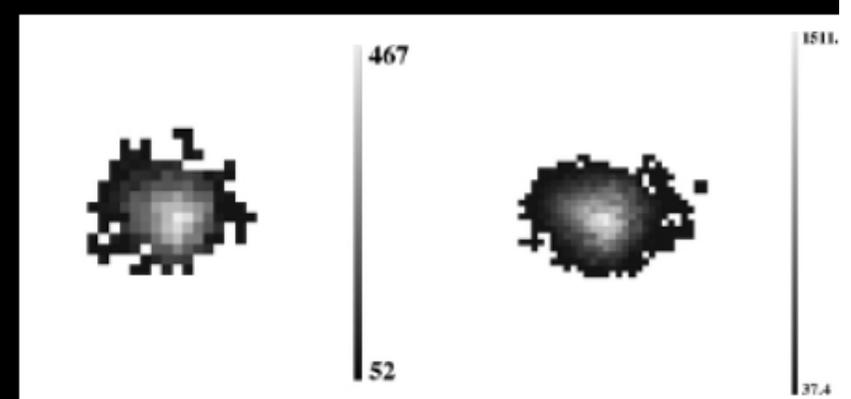
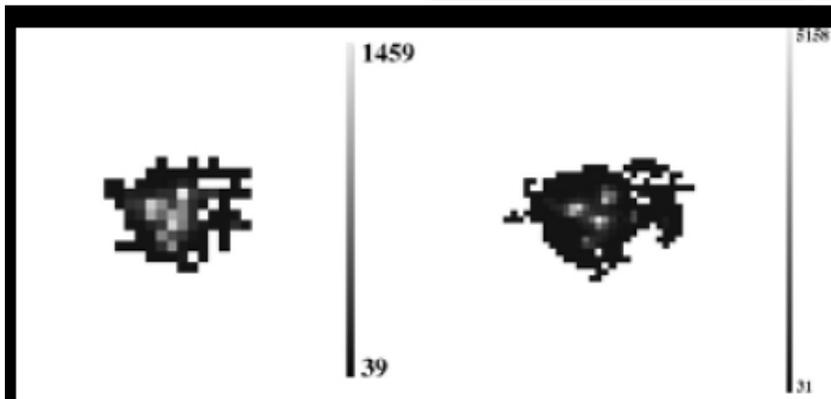
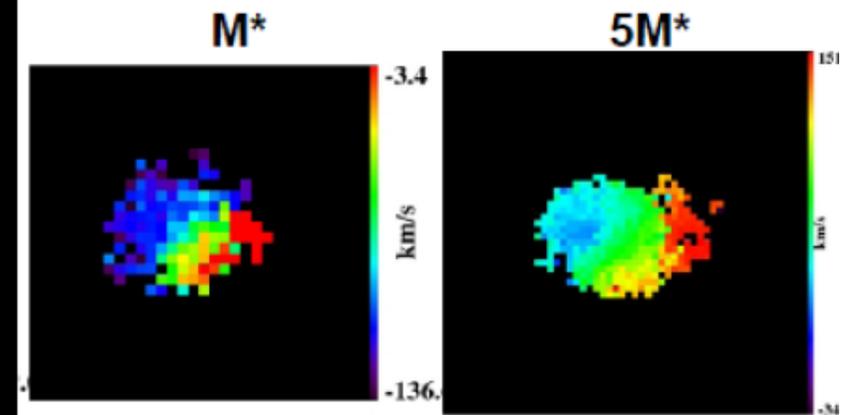
50mas/pix
R=5000



Bournaud+07



GLAO z=4

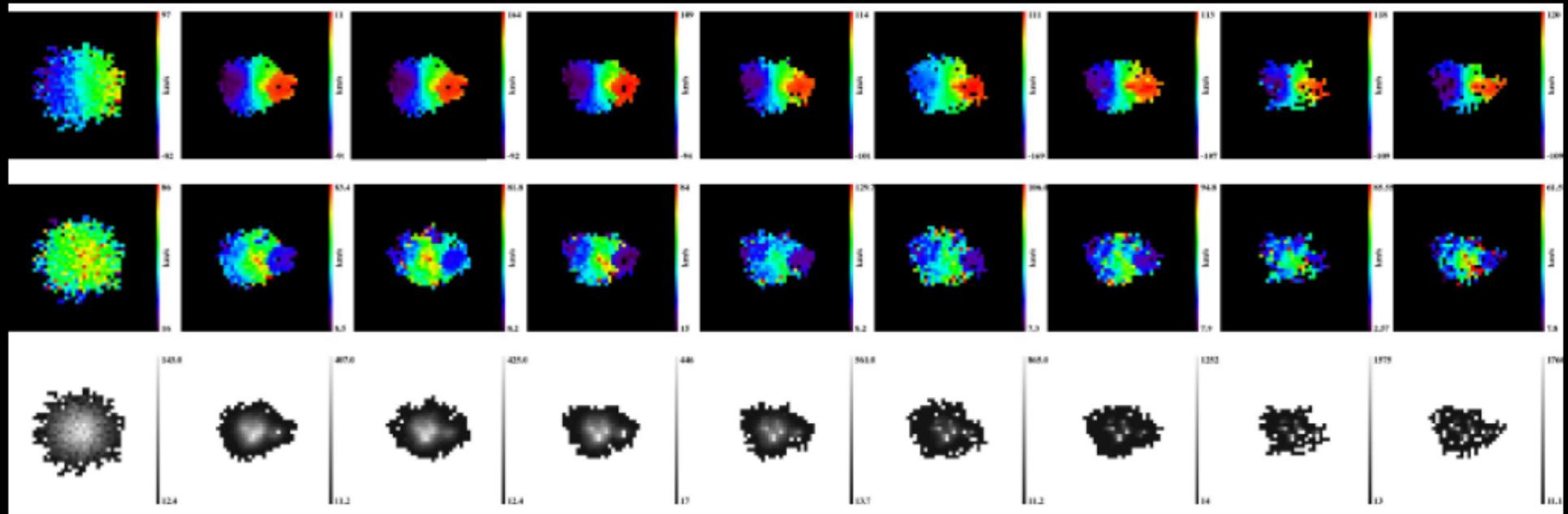


Exemple (2) : EAGLE Phase A

Clumpy Disks

Analogs with a range of EEs ... challenging since we consider a very small distant galaxy at $z=4$ with only 30A EQW (SINFONI data at $z\sim 2$ show $\sim 100A$)

37.5mas/pix (EAGLE baseline)



seeing 15% 61%

At about 20% can start to see clumpy structure, 30% is more robust

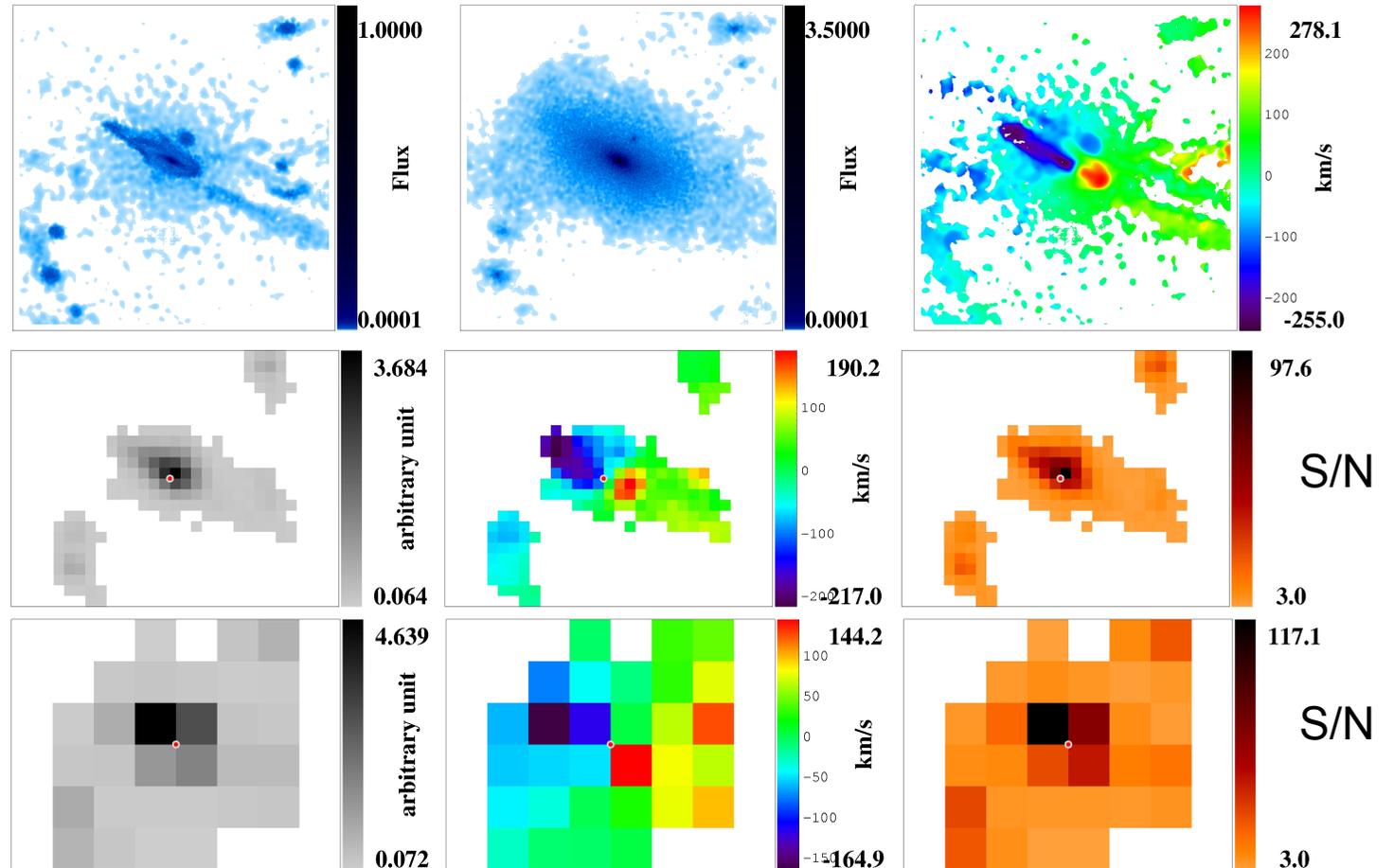
OPTMOS/EVE Phase A: Large IFU: detect and measure extended sources

m(AB)=24 galaxy at z=3 after 10 hours of integration

Top: the hydrodynamic model used as an input to the simulations (shown is the high resolution stellar distribution).

Middle: LI mode simulation. From left to right: flux distribution, velocity field, and the S/N map.

Bottom: same as above but using a spatial binning of 4x4 spaxels.



Delivrables

- Science : D2.1 « Interface definition document »
T0+2m
- Infrastructures : D3.1b COMPASS website
T0+24m
- Management : D1.2 Compass web-based
interface T0+24m
- Instruments designs : D7.2 « ELT-IFU & ELT-
MOS analysis report »

Personnes impliquées au GEPI

GEPI					
GEPI	Puech	Mathieu	Assistant astronome	8	ELT instrumentation specialist, science and instruments design tasks coordinator
GEPI	Hammer	François	Astronome	5	ELT science expertise (PI of OPTIMOS-EVE)
GEPI	Flores	Hector	Assistant astronome	5	ELT instruments expertise (instrument scientist of OPTIMOS-EVE)
GEPI	Disseau	Karen	PhD student	5	ELT instruments preparatory studies. End-to-end simulations for ELT-MOS, link with AO
GEPI	Yang	Yanbin	Post-doc	5	Expertise in instrument simulations and software development
GEPI	Taburet	Sylvestre	Research engineer	3	IT expertise, webmaster of the simulation website
GEPI	Paillous	Michèle	Research engineer	2	Setup and maintenance of hardware
GEPI	Jégouzo	Isabelle	Engineer	4	Software development (website, database)

+ 1 post-doc de 2 ans à recruter